# Siyi Feng

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## **EDUCATION**

Doctor of Natural SciencesSep. 2011-Feb. 2015Universität Heidelberg & Max-Planck-Institut für Astronomie, Heidelberg, Germany•Major: Physics & Astronomy*Thesis for degree:* Chemical and dynamic sub-structures of the high-mass star-forming regions (HMSFRs)

#### Master of Science in Astrophysics

Nanjing University, Jiangsu, China •Major: Astrophysics *Thesis for degree:* Study on the Gamma-Ray Burst (GRB) in Fermi Era

#### **Bachelor of Science in Astronomy**

Nanjing University, Jiangsu, China •Major: Astrophysics *Thesis for degree:* Late Internal Shock Model for the X-ray Flares of the GRB Afterglow

## **RESEARCH INTEREST**

• Star & Planet Formation

• Astrochemistry

Radio Astronomy

## **APPOINTMENTS**

Xiamen University (CN)

-Professor (Nanqiang A) @Department of Astronomy -Associate Professor (Nanqiang B) @Department of Astronomy

#### East Asian Core Observatories Association

-East Asian Core Observatories Association (EACOA) postdoc fellow @NAOJ, NAOC, ASIAA

• Lead the **Multiwavelength line-Imaging survey of the 70**  $\mu$ m-dArk and bright clOuds (MIAO) project. This project carries high-spatial-resolution line-imaging survey towards a sample of 25 high-mass star forming regions. Each region contains a pair of neighbouring 70  $\mu$ m dark/bright clumps, serving as space laboratory for comparative chemical and kinematic study. In particularly, the 70  $\mu$ m dark clump has L/M < 1 L<sub>o</sub>/M<sub>o</sub>. Using JVLA-GBT, IRAM-30 m (120 h, completed, *Feng et al. 2019a, Feng et al. 2020b*), ALMA (90 h, 90% completed), this survey was carried out at 1.3 cm, 3 mm and 1 mm wavelengths, allowing us to characterise the physio-chemical properties of the initial conditions of low-mass and high-mass star formation.

• Lead the multi-scale (1 pc down to 1000 AU) kinematic and chemical study (1 mm/3 mm, NOEMA-IRAM 30 m, SMA-IRAM 30 m) of the high-mass disk candidates NGC 7538 IRS9 (*two 1st-author paper, in prep.*);

• Lead the **effecT of infrared brIght riM bubble (TIME)** line-imaging survey (1mm/3mm, Nobeyama-45 m, IRAM-30 m, Effelsberg-100 m, ALMA) towards the infrared bright rim bubble N2, in understanding the effect of the expanding HII in star formation (*one 1st-author paper, in prep.*);

• Sub-group leader of ALMA large project Fifty AU STudy of the chemistry in the disk/envelope system of Solar-like protostars (FAUST) on the chemistry of low-mass disk candidate GSS30 (*one 1st-author paper, in prep.*);

-Hosts: Dr. K. Tatematsu (2018 in NAOJ); Dr. D. Li (2019 in NAOC); Dr. S. Liu (2020-2022 in ASIAA)

#### Max-Planck-Institute for Extraterrestrial Physics (DE)

-European Research Council (ERC) postdoc fellow @The Center for Astrochemical Studies

• Lead the high spatial resolution line-imaging survey (1mm/2mm/3mm/1.3 cm, NOEMA, SMA, IRAM-30 m, VLA) of sulfur (S-)bearing species and carbon-chains towards the low-mass shocked region L1157-B1 (*Feng et al.* 



Oct. 2021-Present Since Jan. 2023 Oct. 2021–Dec. 2022

Dec. 2017-Sep. 2021

Mar. 2015-Nov.2017

Sep. 2008-Jun. 2011

Sep. 2004-Jun. 2008

2020a, and two 1st-author paper, in prep.); -Director: Prof. Dr. Paola Caselli

### Harvard-Smithsonian Center for Astrophysics (US)

-Visiting Scholar

• Lead the high spatial resolution line-imaging survey (1mm/3mm, IRAM-30 m-SMA) towards four IRDCs, conclude that the initial fragmentation is an non-thermal, non-quiescent process (Feng et al. 2016c);

• Detect two bipolar, high-velocity (up to 40 km s<sup>-1</sup>) outflows (NOEMA) towards a "classic high-mass starless core", conclude that star formation has already begun in this >  $40M_{\odot}$ , <  $10L_{\odot}$  region (*Feng et al. 2016b*). This work has been cited by a graduated-level astrochemistry textbook "Dynamical Astrochemistry" (ISBN-13: 978-1782627760; ISBN-10: 1782627766).

-Host and collaborator: Dr. Qizhou Zhang

#### Max Planck Institute for Astronomy (DE)

-Marie Curie Early Stage Researcher (ESR/PhD) @Planet and Star Formation Group

• Study the chemical structure of the high-mass star-forming regions NGC 7538 S and IRS1 by comparing observations (NOEMA) with gas-grain model fittings, conclude that the fragmentation is hierarchic and that chemical history of the cores collapsed from the same natal cloud is asynchronized (Feng et al. 2016a);

• Study the chemical substructure of the nearest high-mass star-forming region Orion-KL (SMA-IRAM 30 m), conclude that the different spatial distributions of the complex organics indicate their different gas-grain forming paths (Feng et al. 2015). This work has been cited by a graduated-level textbook "Molecular Astrophysics" (ISBN-13: 978-1107169289; ISBN-10: 1107169283).

-Advisor: Prof. Dr. Henrik Beuther

#### Nanjing University (CN)

-Graduated Research Assistant @High Energy Group

• Construct a model of "structured ejecta sweeping up the density-jump medium" to the Gamma-Ray Burst (GRB), which well explains both the flares / bumps on the late afterglows at lower-energy band and the early steep-rising of GeV lightcurve observed by Fermi/ LAT, suggesting the external origin of the GeV photons (Feng et al. 2011); • Simulate the polarization evolution in a relativistic wind bubble, which fits the observed GRB data. -Advisor: Prof. Zigao Dai

#### University of Sheffield (UK)

-Summer Student Researcher @Department of Applied Mathematics

• Revise the models for solar internal f/p/g-modes, including the combined effects from the changes of atmospheric magnetic field, temperature and steady state during a solar cycle (the Best Oral Presentation); -Advisor: Prof. Róbert von Fáy-Siebenbürgen

#### Nanjing University (CN)

-Undergraduate Advanced Project Student Researcher @Department of Astronomy

• Built numerical codes for the dynamic and radiation processes of the "late internal shock" phenomenon towards to GRB, which fits well with the observation of X-ray flares (Excellent Thesis for the Bachelor Degree). -Advisor: Prof. Zigao Dai

## SELECTED OBSERVING PROPOSALS

#### Accepted proposals

• PI, 32.0 h Band 2, JCMT #M22BF001: Dust distribution and chemical segregation in the L1157 molecular outflow • PI, 9.0h (12 m)+47.0h (ACA), ALMA #2019.1.00733.S: The initial gas flow towards extremely young high-mass clumps • PI, 3.5 h (12 m)+18.1 h (ACA), ALMA #2019.1.00408.S: The sequential star formation towards the IR bright rim of an HII bubble • PI, 11.4h (12 m)+90.4h (ACA), ALMA #2018.1.00101.S: The initial gas flow towards extremely young high-mass clumps • PI, 4.4 h (12 m)+20.6 h (ACA), ALMA #2018.1.00215.S: The sequential star formation towards the IR bright rim of an HII bubble • PI, 5.8 h (12 m)+5.6 h (ACA), ALMA #2018.1.00375.S: Sulfur chemistry in the transition zone of low-mass protoplanetary systems • PI, 30 h, JVLA #18A-067: Temperature and density structure of high mass, low luminosity/mass ratio clumps • PI, 18 h, JVLA #18A-068: Detailed physical structure of the protostellar shock region L1157-B1 • PI, 83 h (30 m), IRAM #115-17:

Sep. 2011-Feb. 2015

Sep. 2008-Jun. 2011

Jul.-Aug. 2007

Sep. 2007-Sep. 2008

Dec. 2013-Feb. 2014

Initial star-forming activities towards the high-mass, low luminosity-to-mass ratio clumps • PI, 40 h, Nobevama-45 m #CG171006: Initial star-forming activities towards the high-mass, low luminosity/mass ratio clumps • PI, 25.9 h (30 m), IRAM #009-17: String the beads: Sequential physical and chemical properties of the molecular clumps along the rim of an infrared bubble • PI, 37 h (30 m), IRAM #017-17: Deuterated organics in the earliest phase of high-mass star formation • PI, 44 h (30 m), IRAM #024-17: Sulfur chemistry in the shocked region L1157-B1 & B2 • PI, 9h, APEX #M9505b-99: Deuterated formaldehyde in the earliest phase of high-mass star formation • PI. 20 h. Effelsberg-100 m #105-16: Temperature and density structure of the infrared (IR) bubble rim N2 • PI, 28 h, SMA #2016B-S029: Sulfur and organic chemistry in the shocked region L1157-B1 and B2 • PI, 16 h (NOEMA)+ 1 h (30 m), IRAM #W16AF & 204-16: The main reservoir of sulfur on dust grains in the shocked region L1157-B1 • PI, 15.4 h (30 m), IRAM #036-16: Deuteration in the earliest phase of high-mass star formation • PI, 32 h (NOEMA), IRAM #W14AB: Dynamics and chemistry in the earliest phase of high-mass star formation • PI, 13 h (30 m), IRAM #036-14: Chemistry in the earliest phase of high-mass star formation: SMA & 30m in concert • PI, 20 h, SMA #2013A-S011: Fragmentation and dynamical collapse of high-mass starless gas clumps • PI, 2h (30 m), IRAM #033-12: The chemical substructure of Orion-KL: SMA & 30 m in concert • Co-I, 17.8 h (30 m), IRAM #E01-21: Unveiling the cold Gas Evolution of MaNGA Merging Galaxies (PI: O. Yu, XMU) • Co-I, 25 h (30 m), IRAM #121-20: Filamentary accretion flows in high-mass star-forming molecular clouds (PI: H. Beuther, MPIA) • Co-I, 64 h (NOEMA)+32.5 h (30 m), IRAM #W20AV & 190-20: CORE+: Deuterium and shock chemistry in high-mass star-forming regions (PI: C. Gieser, MPIA) • Co-I, 8h (NOEMA), IRAM #W20BB: Investigating the deuterium fraction in the diffuse and translucent clouds through sensitive absorption observations (PI: G. Luo, NJU) • Co-I, 520 h (band3), JCMT #M20AL021: ALOHA IRDCs: A Lei Of the Habitat and Assembly of Infrared Dark Clouds (PI: D. Li, NAOC) • Co-I, 84 h (band3), JCMT #M20AP020: Dust distribution and chemical segregation in the L1157 molecular outflow (PI: H. Liu, ASIAA) • Co-I, 10.7 h (30 m), IRAM #143-19: Linking density structures and fragmentation in high-mass star formation (PI: H. Beuther, MPIA) • Co-I, 16.3 h, ALMA #2019.1.00280: First detection of magnetic fields in the very central regions of starless dense cores (PI: T. Liu, KASI) • Co-I (the 3rd), 15 h (NOEMA)+13 h (30 m), IRAM #S19-AL & 100-19: Kinematic and chemical signatures during high-mass cloud and star formation (PI: H. Beuther, MPIA) • Co-I, 32 h (band2), JCMT #M18BP002: Are supercritical filaments supported by magnetic fields? (PI: X. Lu, KASI) • Co-I, 28 h (band2), JCMT #M18BP041: Role of magnetic fields in dense core formation in filamentary clouds- II (PI: A. Soam, KASI) • Co-I, 19 h (band2), JCMT #M18BP055: Mapping Magnetic Fields in the Filamentary Cloud NGC 7538 (PI: H. Chen, NTHU) • Co-I, 106.2 h, ALMA #2018.1.01205.L: Fifty AU STudy of the chemistry in the disk/envelope system of Solar-like protostars (PI: S. Yamamoto, UT) • Co-I, 15.7 h, ALMA #2018.1.01449.S: Magnetic fields from infrared dark clouds to hot molecular cores (PI: H. Beuther, MPIA) • Co-I (the 1st), 8 h (NOEMA), IRAM #S18-AN: Linking large and small scales for two high-mass protostars in NGC7538 (PI: J. Mottran, MPIA) • Co-I (the 1st), 37h (NOEMA)+21h (30 m), IRAM #S18-AO, W18-AX, & 222-18, 104-18: Chemical layers of the high-mass disk candidate NGC7538 IRS9 (PI: Y. Wang, MPIA) • Co-I, 53 h (30 m), IRAM #035-18: Early phases of high-mass star formation: kinematics and chemistry in different environments (PI: S. Zhang, OAMP/LAM) • Co-I, 48 h, SMA #2018B-A004: Internal structures of high-mass starless clumps in different environments (PI: J. Yuang, NAOC) • Co-I, 21 h, JVLA #18A-422:

Outflow feedback in early stages of clustered star formation (PI: K. Wang, ESO)

• Co-I, 7.8 h (12 m)+7.4 h (ACA), ALMA #2018.1.00302.S:

Fragmentation and substructure of dense cores close to the onset of star formation in the Orion complex (PI: T. Liu, KASI)

• Co-I, 8.5 h (12 m), ALMA #2017.1.00526.S:

Where and when do low-mass stars form in high-mass protoclusters? (PI: X. Lu, NAOJ)

• Co-I, 9h (12m)+77.6h (ACA), ALMA #2017.1.00687.S:

From filaments to cores: Dynamics in infrared dark clouds (PI: A. Barnes, MPE)

• Co-I, 7.8h (12m)+27.9h (ACA), ALMA #2017.1.00523.S:

Gas accretion onto dense cores from early to late evolutionary phases of massive filamentary clouds (PI: X. Lu, NAOJ)

• Co-I, 11.7 h (ACA), ALMA #2016.2.00058.S:

Physical and chemical properties of cold Orion cores very close to the onset of star formation (PI: K. Tatematsu, NAOJ)

• Co-I (the 1st), 24 h, Nobeyama-45 m #CG161011:

Star formation on the rim of the infrared (IR) bubble N2 (PI: Y. Ao, NAOJ)

• Co-I, 25 h, VLA #16B-259:

Grain growth in the star-forming cluster rho Oph A (PI: A. Coutens, UCL)

• Co-I, 41 h, IRAM-30 m #012-16:

Measuring isotopic ratios in Galactic massive star forming regions with HC3N lines (PI: J. Wang, SHAO) • Co-I, 9.6 h, ALMA #2015.1.00492.S:

Magnetic field structure at the onset of high-mass star (PI: H. Beuther, MPIA)

• Co-I, 618 h, ESO public survey:

Probing the Early Stages of Star Formation: Unravelling the Structure of Planck Cold Clumps Distributed Throughout the Sky (PI: K. Wang, ESO)

• Co-I, 4 h, VLA #15A-115:

Disk and jet formation around the 30Msun protostar NGC7538IRS1 (PI: H. Beuther, MPIA)

• Co-I, 299 h (NOEMA)+8 h (30 m), IRAM large programe #L14AB & 247-13:

Fragmentation and disk formation during high-mass star formation (PI: H. Beuther, MPIA)

• Co-I, 16 h (NOEMA)+2.5 h (30 m), IRAM large programe #W06E & 230-12:

Small-scale fragmentation of genuine high-mass starless cores (HMSCs) (PI: H. Beuther, MPIA)

### **CONFERENCE CONTRIBUTIONS**

• ICE project: The protostellar shocks as chemical laboratories -Invited Talk for 2022 ISM workshop, Nanjing, China	(Aug, 2022)
<ul> <li>Travel through time and space</li> <li>-Invited Public Talk for Xiamen University Malaysia Campus</li> </ul>	(Nov, 2021)
• MIAO: how to draw the profile of initial star-forming regions -Invited Talk for the Cross-Strait Forum on Radio Astronomy 2022 -Talk for Astrochemistry workshop, Zhuhai, China -Invited Talk for NAOJ ALMA-J seminar, Mitaka, Japan -Talk for JingGuangXia workshop, Xiamen, China	(Oct, 2022) (Dec., 2021) (Jan., 2021) (Dec., 2020)
• What can chemistry tell us about the initial conditions and feedbacks of star-formatio -Talk on PSF coffee, Heidelberg, Germany -Talk on Nanjing University Seminar, Nanjing, China -Talk for 2019 ISM workshop, Xinjiang, China -Talk for NAOJ seminar, Mitaka, Japan	n (Oct., 2019) (Sep., 2019) (Jul., 2019) (Jun., 2019)
• Astrochemistry tool: from the star formation to the cradle of life -Talk for UCL seminar, London, UK -Invited Talk for EAO seminar, Hilo, USA -Talk during visit ASIAA, Taipei -Invited Talk during visit SJTU, ZJU, NJU, in China	(Jul., 2018) (May., 2018) (Mar., 2017) (MayOct., 2020)
• Star-forming activities towards extremely cold, young, high-mass star-forming regions -Talk on the "Tracing the flow", Windermere, UK	<b>s</b> (Jul., 2018)
• Sequential physical and chemical properties of the molecular clumps along the rim of -Poster on the "ALMA/Nobeyama/ASTE workshop", Mitaka	an infrared bubble (Dec., 2017)
<ul> <li>Deuteration in the earliest phase of high-mass star formation</li> <li>-Talk on the "Workshop on interstellar matter 2018", Sapporo</li> <li>-Talk on the "MPIA star formation coffee", Heidelberg</li> </ul>	(Nov., 2018) (Jul., 2017)
• G28.34S, a prestellar or protostellar object? -Talk on the "EA-ALMA 2017 meeting", Taipei	(Mar., 2017)
<ul> <li>Outflow detection in a 70 micron dark high-mass core</li> <li>Talk on the "European Week of Astronomy and Space Science 2016", Athens</li> </ul>	(Jul., 2016)
• Chemistry and kinematics in high-mass star-forming regions -Invited Talk during visit NJU, NAOC, KIAA, SAO, China	(Jan, 2016)

• Are the Infrared Dark Clouds Really Quiescent? -Poster on the "From clouds to protoplanetary disks: the astrochemical link", Berlin -Poster on the "Soul of High-Mass Star Formation", Puerto Varas, Chile	(Oct., 2015) (Mar., 2015)
• Complex Organic Molecules in Hot Molecule Cores -Talk on the "Complex Organic Molecules in Space", Pisa -Invited Talk on the "Chemical diagnostics of star and planet formation with Cycle 3 ALMA", MPE	(Mar., 2016) (Jan., 2015)
<ul> <li>Chemical Substructure in High-mass Star-forming Regions</li> <li>-Talk on the "Soul of High-Mass Star Formation", Puerto Varas, Chile</li> <li>-Joint colloquium of MPIA and LSW Talk, MPIA</li> <li>-Radio and Geoastronomy Lunch Talk, Harvard-Smithsonian Center for Astrophysics</li> <li>-Talk on the "Plane &amp; Star Formation Seminar", MPIA</li> </ul>	(Mar., 2015) (Dec., 2014) (Feb., 2014) (Dec., 2013)
<ul> <li>Inferring the Evolutionary Stages of NGC 7538S and NGC 7538 IRS1 from Chemistry</li> <li>Talk on "Chemical Diagnostics in the ALMA/NOEMA Era", Heidelberg</li> <li>Talk on "The Star Formation: Data, Models and Visualization–Harvard-Heidelberg Workshop"</li> <li>Poster on "The Early Phase of Star Formation", Ringberg Castle, Germany</li> <li>Talk on the 223rd American Astronomical Society Meeting, Washington DC, USA</li> <li>Talk on the 17th Annual German Conference of Women in Physics</li> <li>Talk on Conference of "Astrochemistry in the ALMA era", Copenhagen</li> <li>Poster on "High-Mass Star Formation, From Large to Small Scales in the Era of Herschel &amp; ALMA" Lorentz Center, Leiden</li> </ul>	(Jul., 2014) (Jun., 2014) (Jun., 2014) (Jan., 2014) (Nov., 2013) (Jan., 2013) , (Jan., 2013)
<ul> <li>Resolving the Chemical Substructure of Orion-KL</li> <li>-Poster on "Protostar &amp; Planet VI", Heidelberg</li> <li>-Poster on Conference of "Astrochemistry in the ALMA era", Copenhagen</li> <li>-Talk on "2012 MPIA Students Workshop", Bar-sur-Seine</li> </ul>	(Jul., 2013) (Jan., 2013) (Mar., 2012)
• Statistical Characteristics of Interstellar Turbulence, -Talk on the 7th Generation IMPRS Seminar	(Jan., 2012)
<ul> <li>Chemical sub-structure of high-mass star-forming regions,</li> <li>-Poster on Young Astronomers' Meeting (YAM), at Observatoire de Paris</li> </ul>	(Nov., 2011)
<ul> <li>Multiband Fitting to 3 Long GRBs with Fermi/LAT Data: Structured Ejecta Sweeping up a Density-Jump Medium</li> <li>Talk on "Mini Workshop for the Frontier of GRB Research", Nanjing University</li> </ul>	(Nov., 2010)

# SKILLS

Chinese (Mandarin): Mother LanguagePackage (MIRIAD,English: ProfessionalCoding (C, F90, MGerman: BeginnerDigital graphic andJapanese: BeginnerWeb-site developm	IDL, GILDAS, CASA) Itlab, Mathematica, Python, LaTeX) database development ent
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# AWARDS/FELLOWSHIP

2017-2020: 2017-2020: 2015-2017: 2011-2014: 2010: 2008: 2008: 2008: 2005-2008: 2005-2008: 2006:	East Asian Core Observatories Association (EACOA) fellow NAOJ-ALMA fellow (declined) MPE European Research Council (ERC) postdoc fellow Marie Curie Seventh Framework Program Early Stage Researcher (ESR/PhD) Outstanding Contribution To Chinese Astronomical Society Award Graduate with the Highest Honor (top 5%) of Nanjing University Excellent Bachelor Thesis of Nanjing University Undergraduate Student's Outstanding Contribution to Astronomy Outreach Award People's Scholarship for Excellence in Undergraduate Study Outstanding Undergraduate Student Association President in Jiangsu Province of China

# SERVICE

2024/2025 (expected): 2024/2025 (expected):	LOC of "Soul of High-Mass Star Formation" (international conference every 4 years) LOC of "The 2020 East-Asian Young Astronomer Meeting" @Beijing, China
2021–: 2020–2021:	TAC of James Clerk Maxwell Telescope
2014.6:	LOC of "The Star Formation: Data, Models and Visualization-Harvard-Heidelberg Workshop"
2014.6:	LOC of "The Early Phase of Star formation" (EPoS) @Muchen, Germany
2013.7:	LOC of "Protostar & Planets VI" (PPVI) @Heidelberg, Germany
2011-2015:	Guide tours for the planetarium @ Haus de Astronomie, Heidelberg
2008.7:	LOC of 2009 "CAS-IAU Joint Solar Eclipse Meeting" @Suzhou, China
2008.7:	LOC of 2008 "Gamma-Ray Burst Conference" @Nanjing, China